

Users' Requirements for Orbit Stabilization

SPring-8/RIKEN

Tetsuya Ishikawa

Golden Rules

Requests from SR Users are Unlimited.

**Users doubt Light Source before
doubting their Own Equipments.**

**Collaboration between machine people
and users is essential to achieve ultimate
performance.**

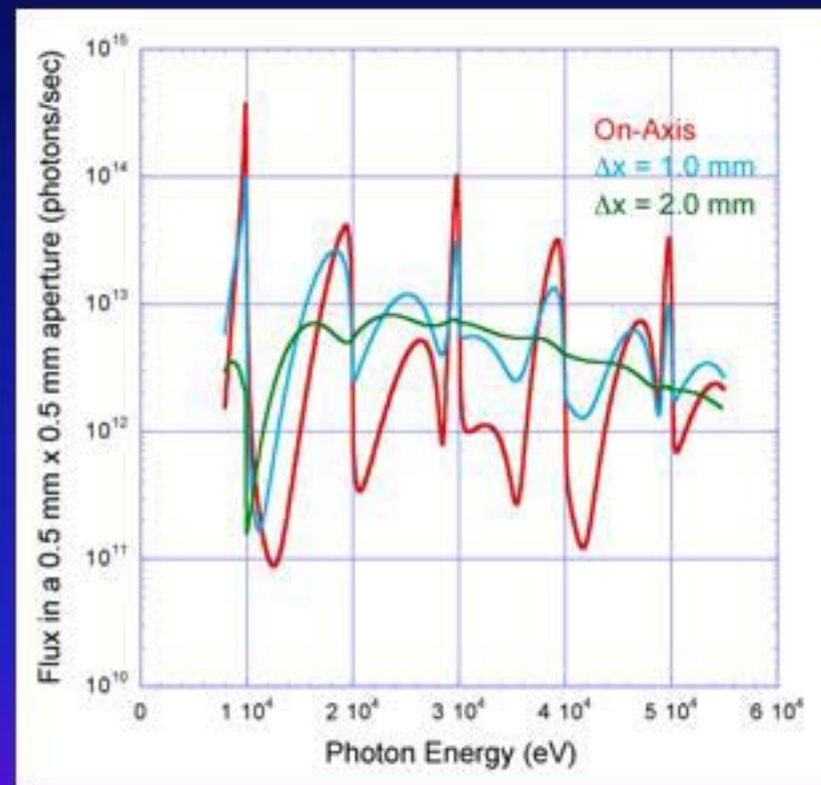
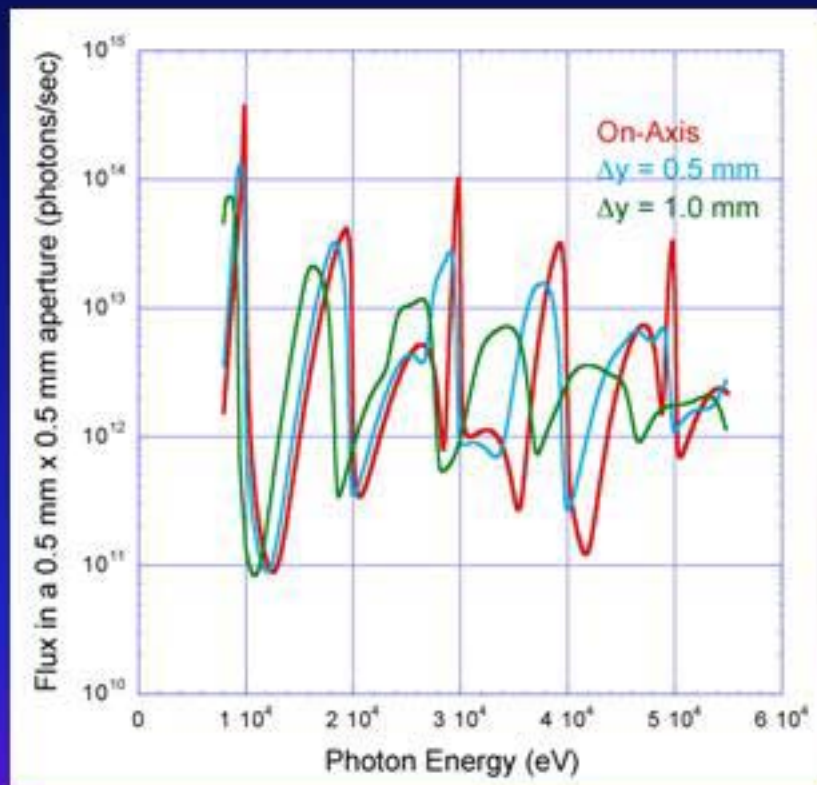
Introduction

- Why the Beam Stability is Important for Users?
- How the Better Stability Improve the Quality of User Data?
- What New Sciences will Come Up with the Further Stability of the Beam?

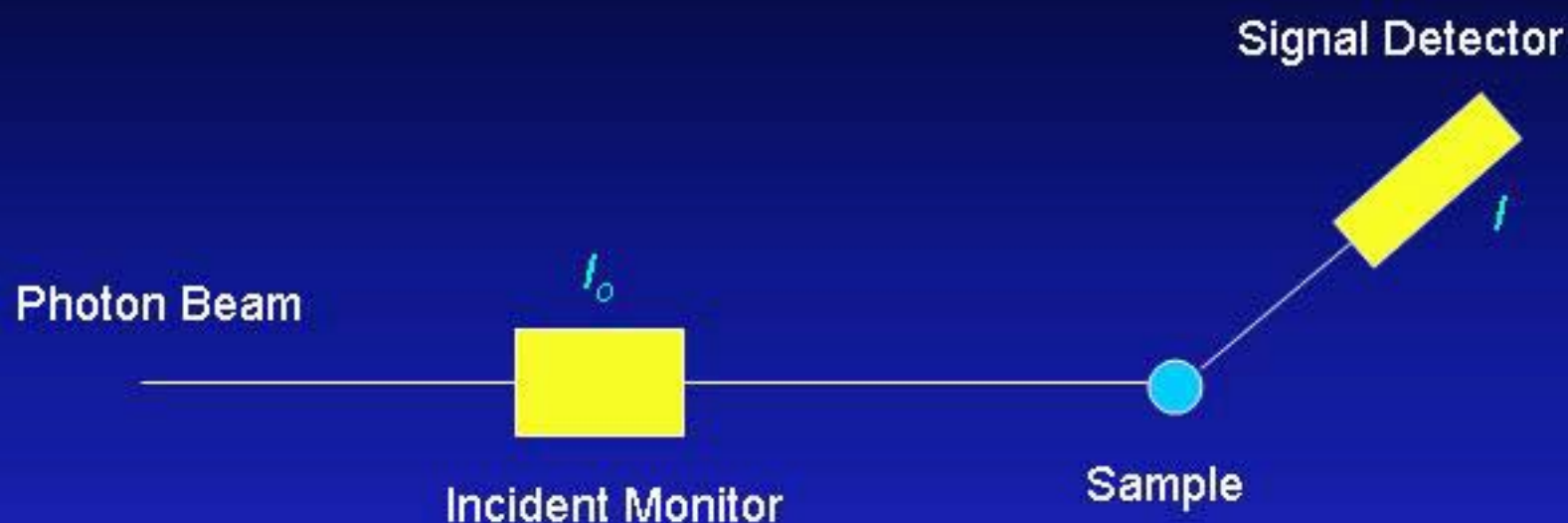
Why the Beam Stability is Important for Users?

- SR Beam is Highly Directional.
- Changes in Source Position will Change the Photon Flux onto Samples.
- Changes in Beam Inclination will Change the Beam Parameters.

Undulator Spectrum



Usual Setup of User Experiments

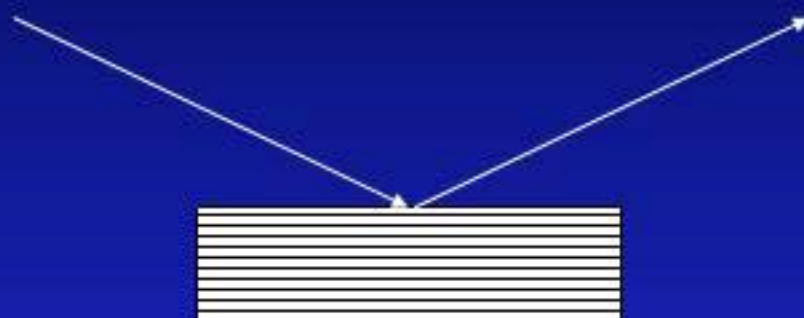


Since storage ring is a time-varying light source, users normalize the signal, I , by incident flux, I_0 .

With fast fluctuation of incident flux, normalization process does not go well.

Spectroscopy: X-ray Monochromatization

Crystal Monochromator



Bragg Condition

$$\lambda = 2d \sin \theta_B$$

λ : wavelength

d : lattice constant

θ_B : glancing angle

Angular Deviation: $\Delta\theta$

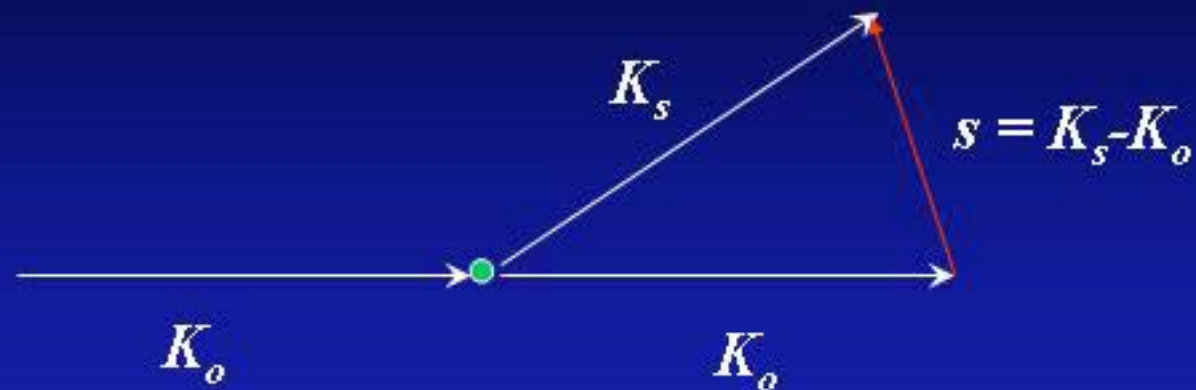
$$\Delta\lambda = 2d \cos \theta_B \Delta\theta$$

Angular Fluctuation



Blurring Fine Structures in
Energy Spectrum

Scattering Measurements

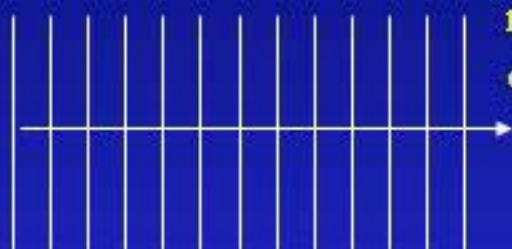


Important Parameter: Scattering Vector, s

$$\text{Angular Fluctuation, } \Delta K_o \quad \longrightarrow \quad \Delta s = - \Delta K_o$$

Scattering of x-rays by a point charge (Thomson Scattering)

$$I_0 = \sqrt{\frac{\epsilon_0}{\mu_0}} E_0^2$$



Point Charge,
mass= m ,
charge= e

$$\mathbf{E}_{\text{radiation}} = \frac{1}{4\pi\epsilon_0 c^2} \frac{e(\ddot{\mathbf{x}} \times \mathbf{r}) \times \mathbf{r}}{r^3}$$

$$I = \sqrt{\frac{\epsilon_0}{\mu_0}} \left(\frac{e^2 \sin^2 \alpha}{4\pi m r c^2} E_0 \right)^2 = r_0^2 \sin^2 \alpha \left(\frac{I_0}{r^2} \right)$$

$$r_0 = \frac{e^2}{4\pi\epsilon_0 m c^2} : \text{Classical Electron Radius}$$

$$= 2.81776 \times 10^{-15} \text{ m}$$

Lorentz Force

$$m\dot{\mathbf{x}} = e(\mathbf{E} + \dot{\mathbf{x}} \times \mathbf{B})$$

$$|\dot{\mathbf{x}}| \ll c \Rightarrow \ddot{\mathbf{x}} = \frac{e}{m} \mathbf{E}$$

Electromagnetic Plane Wave

$$\mathbf{E} = \mathbf{E}_0 \exp[i(\mathbf{K}_0 \cdot \mathbf{r} - \omega t)]$$

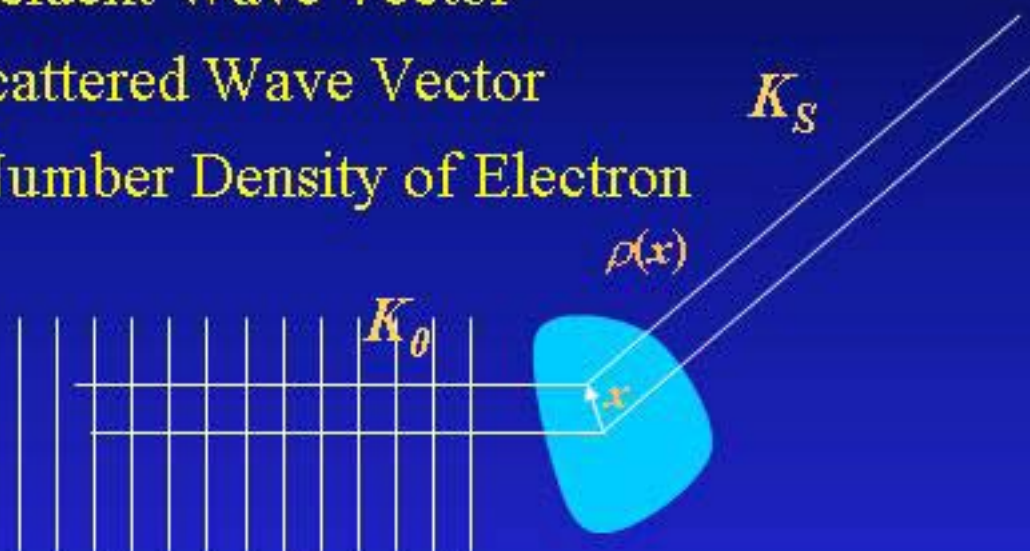
$$\mathbf{B} = \frac{\mathbf{k}}{\omega} \times \mathbf{E}_0 \exp[i(\mathbf{K}_0 \cdot \mathbf{r} - \omega t)] = \frac{\hat{\mathbf{k}}}{c} \times \mathbf{E}$$

Scattering of x-rays by distributed charge

K_θ : Incident Wave Vector

K_s : Scattered Wave Vector

$\rho(\mathbf{x})$: Number Density of Electron



Contribution from a volume element $d^3\mathbf{x}$ at \mathbf{x}

$$dE_{\text{radiation}}^{\text{distribution}} = E_{\text{radiation}}^{\text{point charge}} \rho(\mathbf{x}) \exp[-i(\mathbf{K}_s - \mathbf{K}_o) \cdot \mathbf{x}] d^3\mathbf{x}$$

Scattering of x-rays by distributed charge

$$\mathbf{E}_{radiation}^{distribution} = \mathbf{E}_{radiation}^{point\ charge} \iiint \rho(\mathbf{x}) \exp[-i(\mathbf{K}_s - \mathbf{K}_o) \cdot \mathbf{x}] d^3 \mathbf{x}$$

3D Fourier Transform of Charge Density

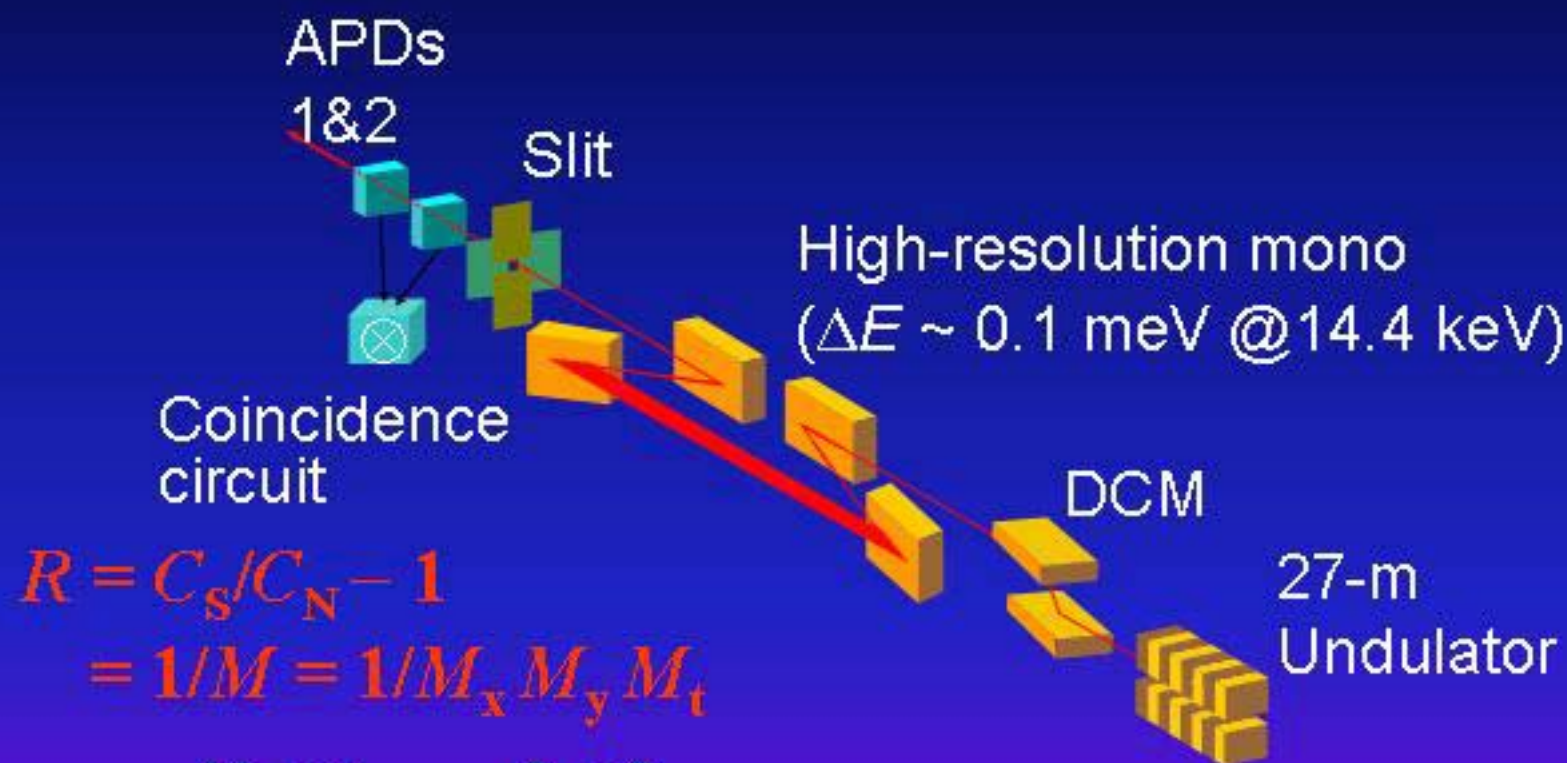
Scattered Intensity

$$\begin{aligned} I &= I^{single} \left| \iiint \rho(\mathbf{x}) \exp[-i(\mathbf{K}_s - \mathbf{K}_o) \cdot \mathbf{x}] d^3 \mathbf{x} \right|^2 \\ &= \frac{r_o^2 \sin^2 \alpha}{r^2} I_o \left| \iiint \rho(\mathbf{x}) \exp[-i(\mathbf{K}_s - \mathbf{K}_o) \cdot \mathbf{x}] d^3 \mathbf{x} \right|^2 \end{aligned}$$

From User Side.....

- **Beam Stability and Optics Stability are Equally Important.**
- **Better Beam Stability Elicits Finer Problems in Optics Instability.**
- **Reversely, Improved Optics Elicits Smaller Beam Instability.**

X-ray Intensity Interferometry (Hambury Brown-Twiss Interferometer)

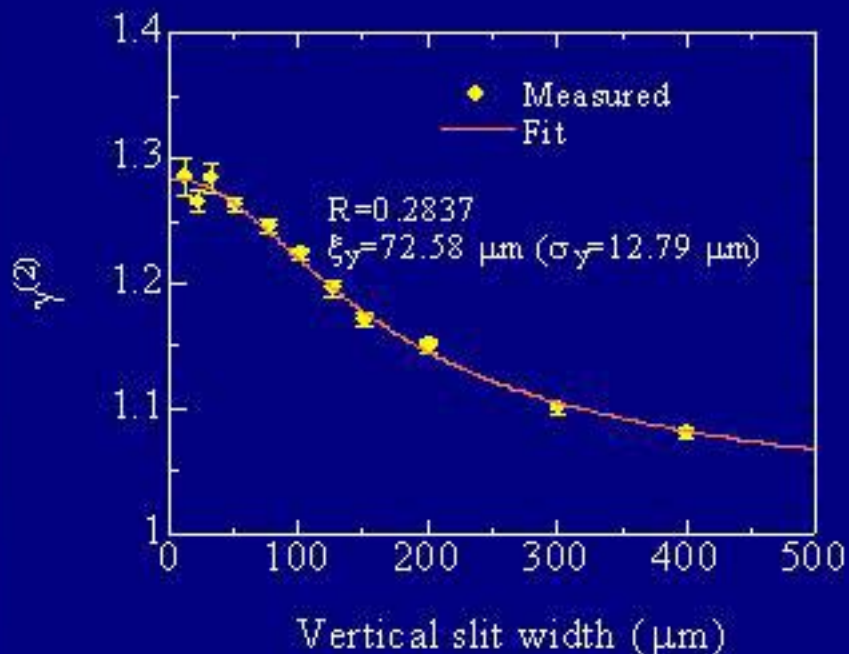


$$R = C_S / C_N - 1$$
$$= 1/M = 1/M_x M_y M_t$$

$$M \rightarrow \infty \quad R \rightarrow 0$$

$$M \rightarrow 1 \quad R \rightarrow 1$$

Source Profile



$\gamma^{(2)} \text{ max} = 1.28$
(@ slit size $10 \times 10 \mu\text{m}^2$)

Coherence length:

$$\xi_y = 72.6 \mu\text{m}$$



Source size

$$\sigma_y = \lambda L / 2\pi \xi_y = 12.8 \mu\text{m}$$

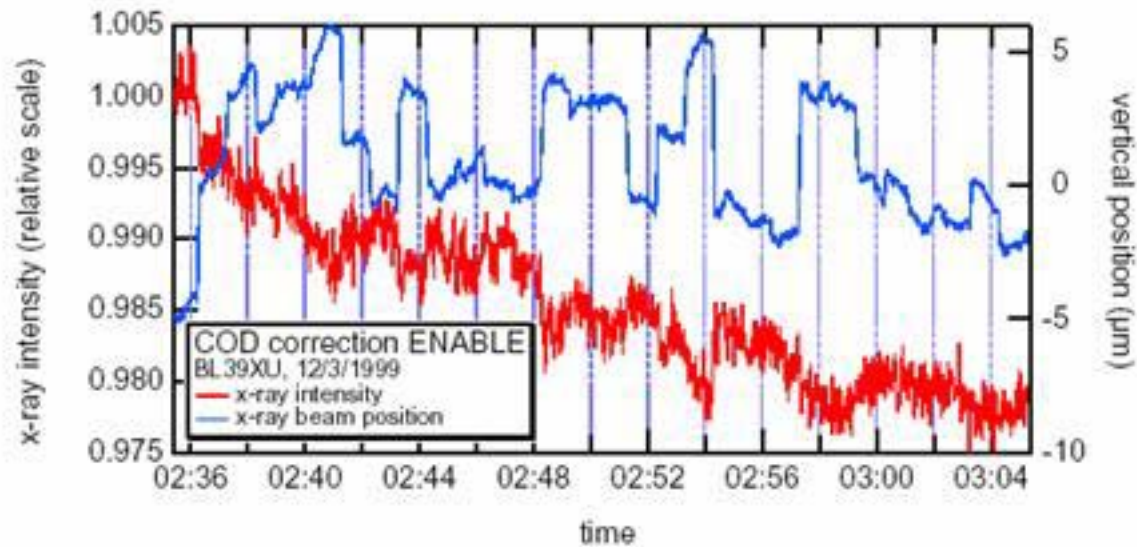
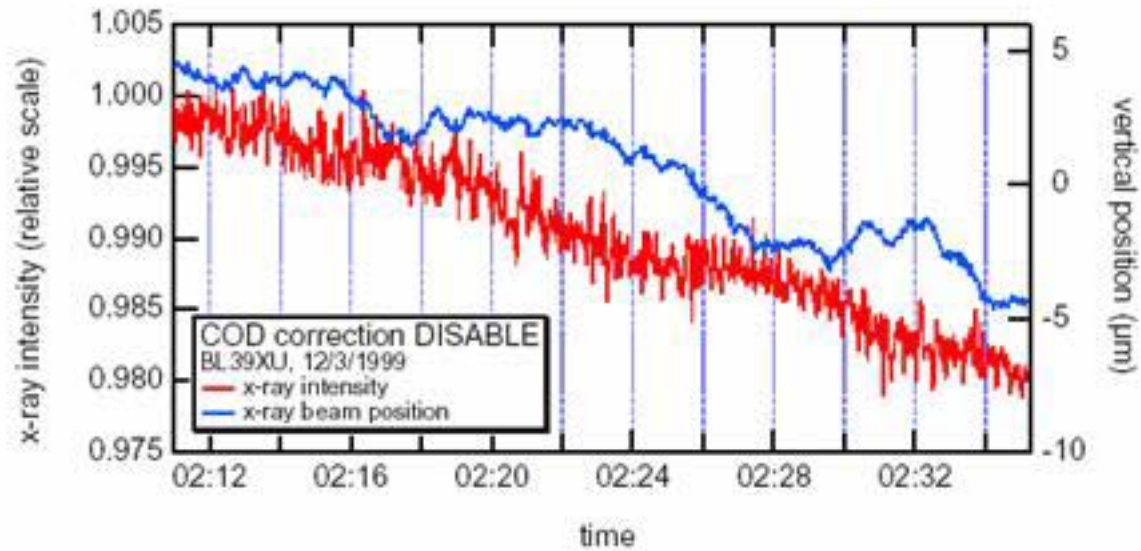
(Accelerator group
measurement:
 $15.9 \pm 1.5 \mu\text{m}$)

Yabashi, Tamazaki & Ishikawa, *PRL* (2004)

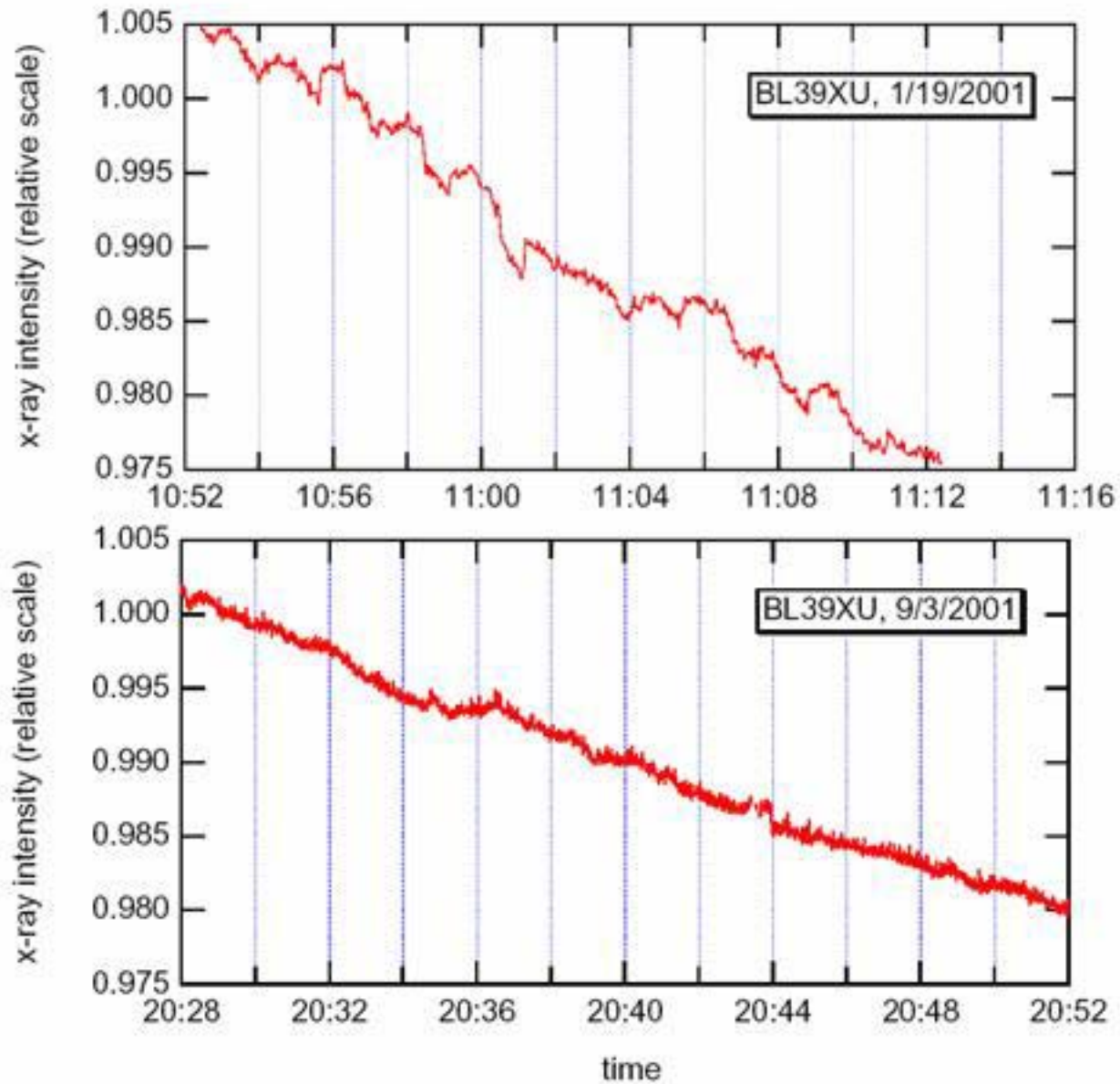
Users may be good monitors of orbit stability.

Collaboration between Machine People and Users is Important.

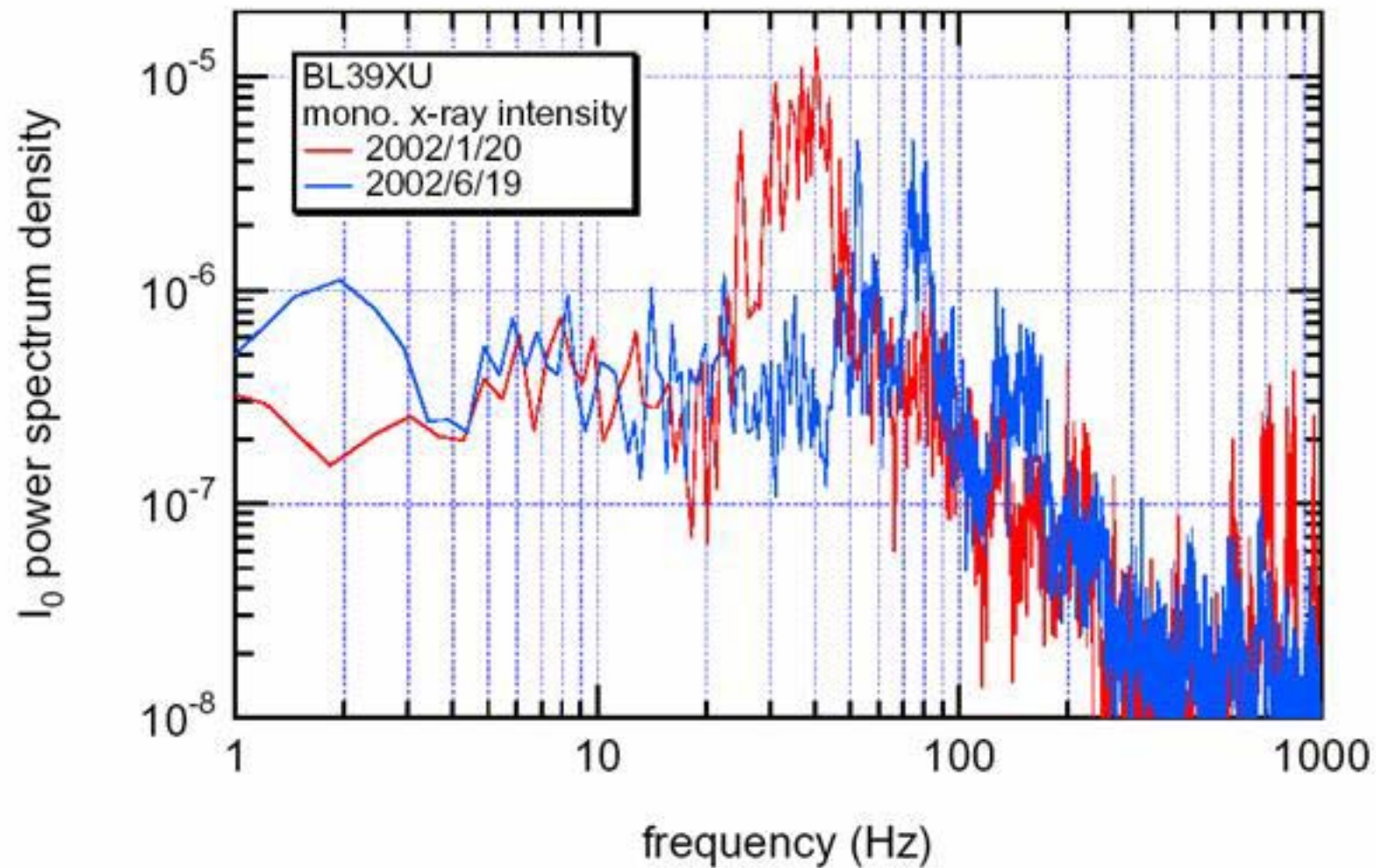
COD Correction: December 1999



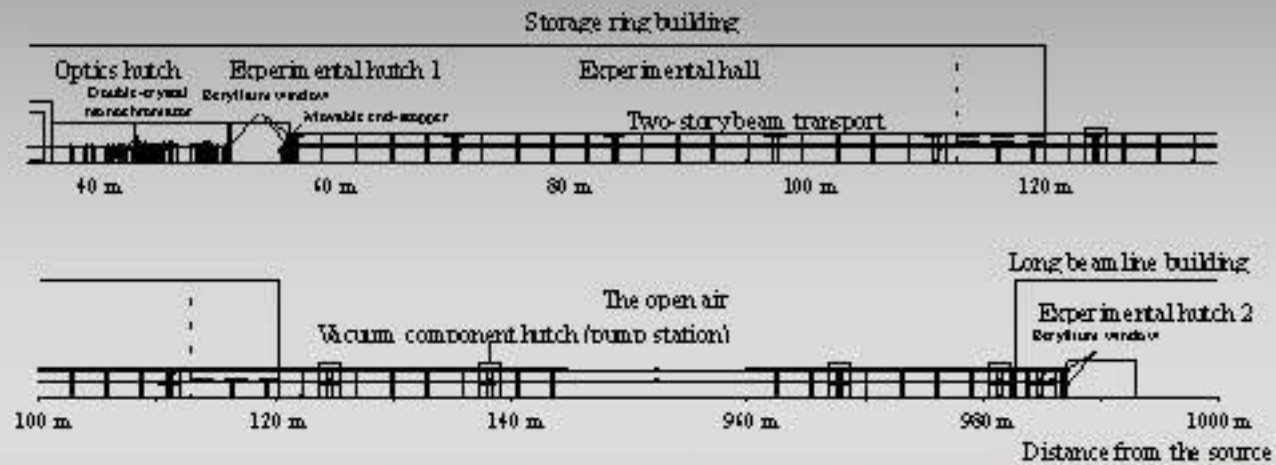
COD Correction: January 2001 vs September 2001



Frequency Power Spectrum: January 2002 vs June 2002



1000 m Beamline, BL29XU



Ishikawa *et al*, *Proc. SPIE* (2001)

What New Sciences will Come Up with the Further Stability of the Beam?

- Phase-Sensitive Imaging (such as holography) demands ultimate overall stability.
- Phase Retrieval of Coherent Scattering

Oversampling Phase Retrieval

What we can measure in diffraction/scattering experiment
= Intensity

All phase information is lost!

Non-Crystalline Charge Distribution:

$$I = \frac{r_o^2 \sin^2 \alpha}{r^2} I_o \left| \iiint \rho(\mathbf{x}) \exp[-i\mathbf{K} \cdot \mathbf{x}] d^3 \mathbf{x} \right|^2 = \sqrt{\frac{\epsilon_o}{\mu_o}} \mathbf{E} \cdot \mathbf{E}^*$$

\mathbf{E}^* : complex conjugate of \mathbf{E}

$$\begin{aligned} \text{If } \mathbf{E} &= \frac{r_o \sin \alpha}{r} \sqrt{I_o} \sqrt{\frac{\mu_o}{\epsilon_o}} \iiint \rho(\mathbf{x}) \exp[-i\mathbf{K} \cdot \mathbf{x}] d^3 \mathbf{x} \\ &= \sqrt{I} \sqrt{\frac{\mu_o}{\epsilon_o}} \exp(i\varphi) \end{aligned}$$

is obtained, we can calculate $\rho(\mathbf{x})$ by Fourier inversion.

Iterative Phase Retrieval

(Jianwei Miao & David Sayre)

X-ray intensity data: Phase Information is Lost!

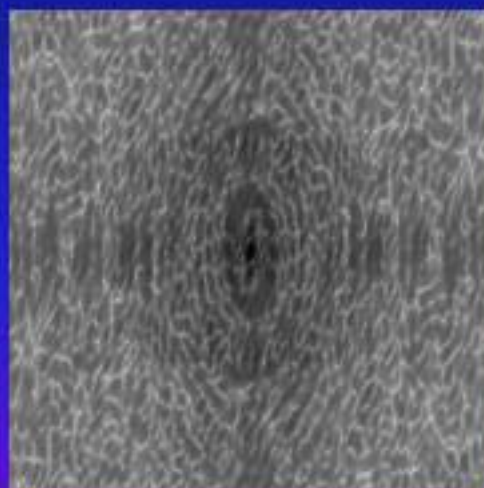
Scattered pattern in Far Field with Coherent Illumination, Phase can be retrieved.

Phase Retrieval → Iterative Algorithm developed by Gerchberg & Saxon, followed by the improvement by Fienup (Opt. Lett. 3 (1978) 27.)



Real Space Image

2002/12/12



Scattered Intensity

Workshop on Beam Orbit Stabilization@SPring-8



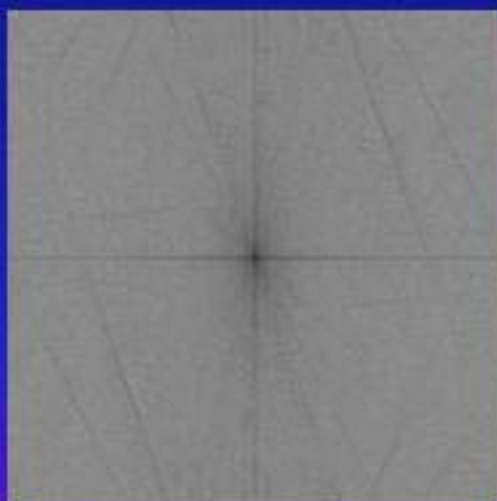
Phase Retrieval

22

Reconstruction of Complex Real Space Images



Real Space



Scattered Intensity



Phase Retrieval



Original Image

**Reconstructed Image
After 5000 iteration**



3D Diffraction Nanoscopy

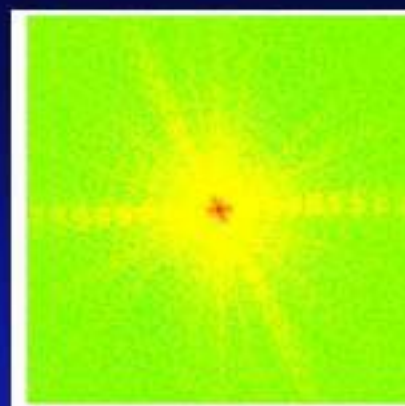
Miao *et al.* PRL (2002)

Two Layer Ni Pattern

8 nm resolution



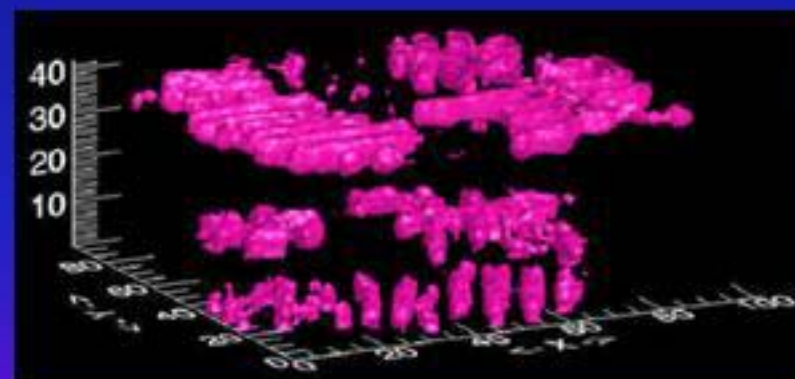
SEM image of Ni pattern on SiN



Coherent Scattering Pattern

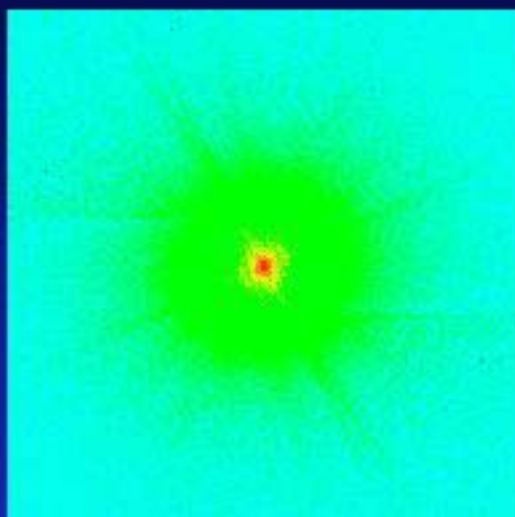


2D Reconstructed Image
(<10 nm resolution)



3D Reconstructed Image (~ 50 nm resolution)

E. Coli Bacteria



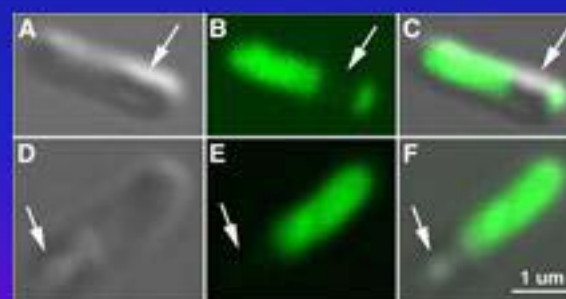
Coherent Scattering
Pattern

25 nm resolution

J. Miao, et al: PNAS, in press



Reconstructed Image



Visible, Confocal Microscopy

Summary

This talk gave addresses to:

- Why the Beam Stability is Important for Users?
- What New Sciences will Come Up with the Further Stability of the Beam?

Acknowledgement

Thank you for your

Toru Hara (Calculation of Undulator Spectra)

Ken Tanasaku and Makina Yabashi (Upgrade of Optics)

Jianwei Li and Y. Ishino, N. Shirai (Coherent Scattering)

Motohiro Suzuki (COD correction effect)

Attention!

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